# Package ‘stars’

July 21, 2022

**Title**  Spatiotemporal Arrays, Raster and Vector Data Cubes

**Version**  0.5-6

**Description**  Reading, manipulating, writing and plotting
spatiotemporal arrays (raster and vector data cubes) in ‘R’, using ‘GDAL’
bindings provided by ‘sf’, and ‘NetCDF’ bindings by ‘ncmeta’ and ‘RNetCDF’.

**License**  Apache License

**URL**  https://r-spatial.github.io/stars/,
https://github.com/r-spatial/stars/

**BugReports**  https://github.com/r-spatial/stars/issues/

**AdditionalRepositories**  http://gis-bigdata.uni-muenster.de/pebesma/

**LazyData**  true

**Depends**  R (>= 3.3.0), abind, sf (>= 1.0-8)

**Imports**  methods, parallel, classInt (>= 0.4-1), lwgeom, rlang, units

**Suggests**  PCICt, RNetCDF (>= 1.8-2), clue, covr, cubelyr, digest, dplyr
(>= 0.7-0), exactextractr, FNN, future.apply, ggforce, ggplot2,
ggthemes, gstat, htrr, jsonlite, knitr, maps, mapdata,
ncdfgeom, ncmeta (>= 0.0.3), pbapply, plm, randomForest,
raster, rgdal, rmarkdown, sp, spacetime, spatstat (>= 2.0-1),
spatstat.geom, starsdata, terra (>= 1.4-22), testthat, tidyrr,
viridis, xts, zoo

**VignetteBuilder**  knitr

**Encoding**  UTF-8

**RoxygenNote**  7.2.0


**NeedsCompilation**  no
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aggregate.stars

spatially or temporally aggregate stars object

Description

spatially or temporally aggregate stars object, returning a data cube with lower spatial or temporal resolution

Usage

## S3 method for class 'stars'
aggregate(
  x,
  by,
  FUN,
  ..., 
  drop = FALSE,
  join = st_intersects,
  as_points = any(st_dimension(by) == 2, na.rm = TRUE),
  rightmost.closed = FALSE,
  left.open = FALSE,
  exact = FALSE
)

Arguments

x

object of class stars with information to be aggregated

by

object of class sf or sfc for spatial aggregation, for temporal aggregation a vector with time values (Date, POSIXct, or PCICt) that is interpreted as a sequence of left-closed, right-open time intervals or a string like "months", "5 days" or the like (see cut.POSIXt); if by is an object of class stars, it is converted to sfc by st_as_sfc(by, as_points = FALSE) thus ignoring its time component. Note: each pixel is assigned to only a single group (in the order the groups occur) so non-overlapping spatial features and temporal windows are recommended.

FUN

aggregation function, such as mean

... arguments passed on to FUN, such as na.rm=TRUE
drop logical; ignored
join function; function used to find matches of x to by
as_points see st_as_sf: shall raster pixels be taken as points, or small square polygons?
rightmost.closed see findInterval
left.open logical; used for time intervals, see findInterval and cut.POSIXt
exact logical; if TRUE, use coverage_fraction to compute exact overlap fractions of polygons with raster cells

Examples

# aggregate time dimension in format Date
tif = system.file("tif/L7_ETMs.tif", package = "stars")
t1 = as.Date("2018-07-31")
x = read_stars(c(tif, tif, tif, tif), along = list(time = c(t1, t1+1, t1+2, t1+3)))[,1:30,1:30]
st_get_dimension_values(x, "time")
x_agg_time = aggregate(x, by = t1 + c(0, 2, 4), FUN = max)

# aggregate time dimension in format Date - interval
by_t = "2 days"
x_agg_time2 = aggregate(x, by = by_t, FUN = max)
st_get_dimension_values(x_agg_time2, "time")
#TBD:
#x_agg_time - x_agg_time2

# aggregate time dimension in format POSIXct
x = st_set_dimensions(x, 4, values = as.POSIXct(c("2018-07-31",
"2018-08-01",
"2018-08-02",
"2018-08-03"),
names = "time")
by_t = as.POSIXct(c("2018-07-31", "2018-08-02"))
x_agg posix = aggregate(x, by = by_t, FUN = max)
st_get_dimension_values(x_agg posix, "time")
#TBD:
# x_agg_time - x_agg posix
aggregate(x, "2 days", mean)
if (require(ncmeta, quietly = TRUE)) {
  # Spatial aggregation, see https://github.com/r-spatial-stars/issues/299
  prec_file = system.file("nc/test_stageiv_xyt.nc", package = "stars")
  prec = read_ncdf(prec_file, curvilinear = c("lon", "lat"))
  prec_slice = dplyr::slice(prec, index = 17, along = "time")
  nc = sf::read_sf(system.file("gpkg/nc.gpkg", package = "sf"), "nc.gpkg")
  nc = st_transform(nc, st_crs(prec_slice))
  agg = aggregate(prec_slice, st_geometry(nc), mean)
  plot(agg)
}

# example of using a function for "by": aggregate by month-of-year
d = c(10, 10, 150)
a = array(rnorm(prod(d)), d)  # pure noise
times = Sys.Date() + seq(1, 2000, length.out = d[3])
m = as.numeric(format(times, "%m"))
signal = rep(sin(m / 12 * pi), each = prod(d[1:2])) # yearly period
s = (st_as_stars(a) + signal) %>%
  st_set_dimensions(3, values = times)
f = function(x, format = "%B") {
  months = format(as.Date(paste0("01-", 1:12, "-1970")), format)
  factor(format(x, format), levels = months)
}
agg = aggregate(s, f, mean)
plot(agg)

---

**as**: Coerce stars object into a Raster raster or brick

**Description**

Coerce stars object into a Raster raster or brick

Coerce stars object into a terra SpatRaster

**Arguments**

from: object to coerce

**Details**

If the stars object has more than three dimensions, all dimensions higher than the third will be collapsed into the third dimensions. If the stars object has only an x/y raster but multiple attributes, these are merged first, then put in a raster brick.

If the stars object has more than three dimensions, all dimensions higher than the third will be collapsed into the third dimensions. If the stars object has only an x/y raster but multiple attributes, these are merged first, then put in a SpatRaster.

**Value**

RasterLayer or RasterBrick

SpatRaster
bcsd_obs  

*Monthly Gridded Meteorological Observations*

**Description**

These are the monthly observational data used for BCSD downscaling. See: http://gdo-dcp.uclnl.org/downscaled_cmip_projections/dcpInterface.html#About for more information."; "Atmospheric Temperature, Air Temperature Atmosphere, Precipitation, Rain, Maximum Daily Temperature, Minimum Daily Temperature";

**Usage**

bcsd_obs

**Format**

An object of class stars_proxy (inherits from stars) of dimension 81 x 33 x 12.

c.stars  

*combine multiple stars objects, or combine multiple attributes in a single stars object into a single array*

**Description**

combine multiple stars objects, or combine multiple attributes in a single stars object into a single array

**Usage**

```r
## S3 method for class 'stars'
c(...,
     along = NA_integer_,
     try_hard = FALSE,
     nms = names(list(...)),
     tolerance = sqrt(.Machine$double.eps)
)

## S3 method for class 'stars_proxy'
c(...,
     along = NA_integer_,
     along_crs = FALSE,
     try_hard = FALSE,
     nms = names(list(...)),
     tolerance = sqrt(.Machine$double.eps)
)
```
Arguments

... object(s) of class star: in case of multiple arguments, these are combined into a single stars object, in case of a single argument, its attributes are combined into a single attribute. In case of multiple objects, all objects should have the same dimensionality.

along integer; see read_stars

try_hard logical; if TRUE and some arrays have different dimensions, combine those that dimensions matching to the first array

tolerance numeric; values used in all.equal to compare dimension values combine those that dimensions matching to the first array

along_crs logical; if TRUE, combine arrays along a CRS dimension

Details

An error is raised when attempting to combine arrays with different measurement units into a single array. If this was inteneded, drop_units can be used to remove units of a stars object before merging.

Value

a single stars object with merged (binded) arrays.

Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
(new = c(x, x))
c(new) # collapses two arrays into one with an additional dimension
c(x, x, along = 3)

contour.stars plot contours of a stars object

Description

plot contours of a stars object

Usage

## S3 method for class 'stars'
contour(x, ...)

Arguments

x object of class stars

... other parameters passed on to contour
Details

this uses the R internal contour algorithm, which (by default) plots contours; st_contour uses the GDAL contour algorithm that returns contours as simple features.

Examples

d = st_dimensions(x = 1:ncol(volcano), y = 1:nrow(volcano))
r = st_as_stars(t(volcano))
r = st_set_dimensions(r, 1, offset = 0, delta = 1)
r = st_set_dimensions(r, 2, offset = 0, delta = -1)
plot(r, reset = FALSE)
contour(r, add = TRUE)

cut_stars     cut methods for stars objects

Description

cut methods for stars objects

Usage

## S3 method for class 'array'
cut(x, breaks, ...)

## S3 method for class 'matrix'
cut(x, breaks, ...)

## S3 method for class 'stars'
cut(x, breaks, ...)

Arguments

x          see cut
breaks     see cut
...         see cut

Details

R’s factor only works for vectors, not for arrays or matrices. This is a work-around (or hack?) to keep the factor levels generated by cut and use them in plots.

Value

an array or matrix with a levels attribute; see details
Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
cut(x, c(0, 50, 100, 255))
cut(x[,,1], c(0, 50, 100, 255))
plot(cut(x[,,1], c(0, 50, 100, 255)))
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x1 = read_stars(tif)
(x1_cut = cut(x1, breaks = c(0, 50, 100, Inf))) # shows factor in summary
plot(x1_cut[,,c(3,6)]) # propagates through [ and plot

---

dplyr verbs for stars objects

dplyr verbs for stars objects; package dplyr needs to be loaded before these methods can be used for stars objects.

Usage

  filter.stars(.data, ...)
  filter.stars_proxy(.data, ...)
  mutate.stars(.data, ...)
  mutate.stars_proxy(.data, ...)
  transmute.stars(.data, ...)
  transmute.stars_proxy(.data, ...)
  select.stars(.data, ...)
  select.stars_proxy(.data, ...)
  rename.stars(.data, ...)
  rename.stars_proxy(.data, ...)
  pull.stars(.data, var = -1)
  pull.stars_proxy(.data, ...)
  as.tbl_cube.stars(x, ...)

Description

dplyr verbs for stars objects; package dplyr needs to be loaded before these methods can be used for stars objects.
slice.stars(.data, along, index, ..., drop = length(index) == 1)
slice.stars_proxy(.data, along, index, ...)
replace_na.stars(data, replace, ...)
replace_na.stars_proxy(data, ...)

Arguments

.data object of class stars
...
see filter
var see pull
x object of class stars
along name or index of dimension to which the slice should be applied
index integer value(s) for this index
drop logical; drop dimensions that only have a single index?
data data set to work on
replace see replace_na: list with variable=value pairs, where value is the replacement value for NA's

Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x1 = read_stars(tif)
if (require(dplyr, quietly = TRUE)) {
x1 %>% slice("band", 2:3)
x1 %>% slice("x", 50:100)
}

---

geom_stars ggplot geom for stars objects

Description

ggplot geom for stars objects

Usage

geom_stars(
mapping = NULL,
data = NULL,
..., downsampling = 0,
sf = FALSE,
na.action = na.pass

theme_stars(...)

Arguments

mapping see geom_raster

data see geom_raster
...

downsample downsampling rate: e.g. 3 keeps rows and cols 1, 4, 7, 10 etc.; a value of 0 does not downsample; can be specified for each dimension, e.g. c(5,5,0) to downsample the first two dimensions but not the third.

sf logical; if TRUE rasters will be converted to polygons and plotted using geom_sf.

na.action function; if NA values need to be removed before plotting use the value na.omit here (only applies to objects with raster dimensions)

Details

geom_stars returns (a call to) either geom_raster, geom_tile, or geom_sf, depending on the raster or vector geometry; for the first to, an aes call is constructed with the raster dimension names and the first array as fill variable. Further calls to coord_equal and facet_wrap are needed to control aspect ratio and the layers to be plotted; see examples. If a stars array contains hex color values, and no fill parameter is given, the color values are used as fill color; see the example below.

Examples

system.file("tif/L7_ETMs.tif", package = "stars") %>% read_stars() -> x
if (require(ggplot2, quietly = TRUE)) {
  ggplot() + geom_stars(data = x) +
  coord_equal() +
  facet_wrap(~band) +
  theme_void() +
  scale_x_discrete(expand=c(0,0))+
  scale_y_discrete(expand=c(0,0))
  # plot rgb composite:
  st_as_stars(L7_ETMs)[,,1:3] |> st_rgb() -> x # x contains colors as pixel values
  ggplot() + geom_stars(data = x)
}

L7_ETMs

Landsat-7 bands for a selected region around Olinda, BR
Description

Probably containing the six 30 m bands:

- Band 1 Visible (0.45 - 0.52 µm) 30 m
- Band 2 Visible (0.52 - 0.60 µm) 30 m
- Band 3 Visible (0.63 - 0.69 µm) 30 m
- Band 4 Near-Infrared (0.77 - 0.90 µm) 30 m
- Band 5 Short-wave Infrared (1.55 - 1.75 µm) 30 m
- Band 7 Mid-Infrared (2.08 - 2.35 µm) 30 m

Usage

L7_ETMs

Format

An object of class stars_proxy (inherits from stars) of dimension 349 x 352 x 6.

make_intervals

create an intervals object

Description

create an intervals object, assuming left-closed and right-open intervals

Usage

make_intervals(start, end)

Arguments

start vector with start values, or 2-column matrix with start and end values in column 1 and 2, respectively
end vector with end values
merge

merge or split stars object

Description
merge attributes into a dimension, or split a dimension over attributes

Usage
## S3 method for class 'stars'
split(x, f = length(dim(x)), drop = TRUE, ...)

## S3 method for class 'stars'
merge(x, y, ..., name = "attributes")

Arguments

x: object of class stars
f: the name or index of the dimension to split; by default the last dimension
drop: ignored
...: if defined, the first unnamed argument is used for dimension values, if not defined, attribute names are used for dimension values
y: needs to be missing
name: name for the new dimension

Details
split.stars works on the first attribute, and will give an error when more than one attribute is present

Value
merge merges attributes of a stars object into a new dimension; split splits a dimension over attributes

ops_stars
S3 Ops Group Generic Functions for stars objects

Description
Ops functions for stars objects, including comparison, product and divide, add, subtract
Usage

```r
## S3 method for class 'stars'
Ops(e1, e2)

## S3 method for class 'stars'
Math(x, ...)

## S3 method for class 'stars_proxy'
Ops(e1, e2)

## S3 method for class 'stars_proxy'
Math(x, ...)
```

Arguments

- `e1`: object of class `stars`
- `e2`: object of class `stars`
- `x`: object of class `stars`
- `...`: parameters passed on to the Math functions

Details

If `e1` or `e2` is a numeric vector, or `e2` has less or smaller dimensions than `e1`, then `e2` is recycled such that it fits `e1`, using usual R array recycling rules. The user needs to make sure this is sensible; it may be needed to use `aperm` to permutate dimensions first.

Value

object of class `stars`

Examples

```r
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
x  
x * x
x / x
x + x
x + 10
all.equal(x * 10, 10 * x)
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
a = sqrt(x)
b = log(x, base = 10)
```
Description

plot stars object, with subplots for each level of first non-spatial dimension, and customization of legend key

Usage

```r
## S3 method for class 'stars'
plot(
x, y, ...
join_zlim = TRUE,
main = make_label(x, 1),
axes = FALSE,
downsamp = TRUE,
nbreaks = 11,
breaks = "quantile",
col = grey(1:(nb - 1)/nb),
key.pos = get_key_pos(x, ...),
key.width = lcm(1.8),
key.length = 0.618,
reset = TRUE,
box_col = grey(0.8),
center_time = FALSE,
hook = NULL,
mfrow = NULL
)
```

```r
## S3 method for class 'stars'
image(
x, ...
band = 1,
attr = 1,
asp = NULL,
rgb = NULL,
maxColorValue = ifelse(inherits(rgb, "data.frame"), 255, max(x[[attr]], na.rm = TRUE)),
xlab = if (!axes) "" else names(d)[1],
ylab = if (!axes) "" else names(d)[2],
xlim = st_bbox(extent)$xlim,
ylim = st_bbox(extent)$ylim,
```
text_values = FALSE,
  text_color = "black",
  axes = FALSE,
  interpolate = FALSE,
  as_points = FALSE,
  key.pos = NULL,
  logz = FALSE,
  key.width = lcm(1.8),
  key.length = 0.618,
  add.geom = NULL,
  border = NA,
  useRaster = isTRUE(dev.capabilities()$rasterImage == "yes"),
  extent = x
)

## S3 method for class 'nc_proxy'
plot(x, y, ..., downsample = get_downsample(dim(x)), max_times = 16)

## S3 method for class 'stars_proxy'
plot(x, y, ..., downsample = get_downsample(dim(x)))

Arguments

x object of class stars

y ignored

... further arguments: for plot, passed on to image.stars; for image, passed on to image.default or rasterImage.

join_zlim logical; if TRUE, compute a single, joint zlim (color scale) for all subplots from x

main character; subplot title prefix; use "" to get only time, use NULL to suppress subplot titles

axes logical; should axes and box be added to the plot?

downsample logical or numeric; if TRUE will try to plot not many more pixels than actually are visible, if FALSE, no downsampling takes place, if numeric, the number of pixels/lines/bands etc that will be skipped; see Details.

nbreaks number of color breaks; should be one more than number of colors. If missing and col is specified, it is derived from that.

breaks actual color breaks, or a method name used for classIntervals.

col colors to use for grid cells

key.pos integer; side to plot a color key: 1 bottom, 2 left, 3 top, 4 right; set to NULL to omit key. Ignored if multiple columns are plotted in a single function call. Default depends on plot size, map aspect, and, if set, parameter asp.

key.width amount of space reserved for width of the key (labels); relative or absolute (using lcm)

key.length amount of space reserved for length of the key (labels); relative or absolute (using lcm)
reset logical; if FALSE, keep the plot in a mode that allows adding further map elements; if TRUE restore original mode after plotting

box_col color for box around sub-plots; use 0 to suppress plotting of boxes around sub-plots.

center_time logical; if TRUE, sub-plot titles will show the center of time intervals, otherwise their start

hook NULL or function; hook function that will be called on every sub-plot.

mfrow length-2 integer vector with nrow, ncolumns of a composite plot, to override the default layout

band integer; which band (dimension) to plot

attr integer; which attribute to plot

asp numeric; aspect ratio of image

rgb integer; specify three bands to form an rgb composite. Experimental: rgb color table; see Details.

maxColorValue numeric; passed on to rgb

xlab character; x axis label

ylab character; y axis label

xlim x axis limits

ylim y axis limits

text_values logical; print values as text on image?

text_color character; color for printed text values

interpolate logical; when using rasterImage (rgb), should pixels be interpolated?

as_points logical; for curvilinear or sheared grids: parameter passed on to st_as_sf, determining whether raster cells will be plotted as symbols (fast, approximate) or small polygons (slow, exact)

logz logical; if TRUE, use log10-scale for the attribute variable. In that case, breaks and at need to be given as log10-values; see examples.

add.geom object of class sfc, or list with arguments to plot, that will be added to an image or sub-image

border color used for cell borders (only in case x is a curvilinear or rotated/sheared grid)

useRaster logical; use the rasterImage capabilities of the graphics device?

extent object which has a st_bbox method; sets the plotting extent

max_times integer; maximum number of time steps to attempt to plot.

Details

Downsampling: a value for downsample of 0: no downsampling, 1: after every dimension value (pixel/line/band), one value is skipped (half of the original resolution), 2: after every dimension value, 2 values are skipped (one third of the original resolution), etc.

To remove unused classes in a categorical raster, use the droplevels function.

use of an rgb color table is experimental; see https://github.com/r-spatial/mapview/issues/208
when plotting a subsetted \texttt{stars\_proxy} object, the default value for argument \texttt{downsample} will not be computed correctly, and has to be set manually.

Examples

\begin{verbatim}
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
image(x, col = grey((3:9)/10))
image(x, rgb = c(1,3,5)) # rgb composite
\end{verbatim}

\texttt{predict.stars} \hspace{1cm} \textit{Predict values, given a model object, for a \texttt{stars} or \texttt{stars\_proxy} object}

Description

Predict values, given a model object, for a \texttt{stars} or \texttt{stars\_proxy} object

Usage

\begin{verbatim}
## S3 method for class 'stars'
predict(object, model, ..., drop_dimensions = FALSE)

## S3 method for class 'stars\_proxy'
predict(object, model, ...)
\end{verbatim}

Arguments

- \texttt{object} \hspace{1cm} object of class \texttt{`stars`}
- \texttt{model} \hspace{1cm} model object of a class that has a predict method; check with \texttt{`methods(class = class(object))`}
- ... \hspace{1cm} arguments passed on to this predict method
- \texttt{drop_dimensions} \hspace{1cm} logical; if \texttt{`TRUE`}, remove dimensions (coordinates etc) from \texttt{`data.frame`} with predictors

Details

separate predictors in object need to be separate attributes in \texttt{object}; in case they are e.g. in a band dimension, use \texttt{`split(object)`}
Read data using GDAL’s multidimensional array API (experimental)

Usage

read_mdim(
  x,
  variable = character(0),
  ...,  
  options = character(0),
  raster = NULL,
  offset = integer(0),
  count = integer(0),
  step = integer(0),
  proxy = FALSE,
  debug = FALSE
)

Arguments

  x          data source name
  variable   name of the array to be read
  ...        ignored
  options    array opening options
  raster     names of the raster variables (default: first two)
  offset     integer; offset for each dimension (pixels) of sub-array to read (default: 0,0,0,...) (requires sf >= 1.0-9)
  count      integer; size for each dimension (pixels) of sub-array to read (default: read all) (requires sf >= 1.0-9)
  step       integer; step size for each dimension (pixels) of sub-array to read (requires sf >= 1.0-9)
  proxy      logical; return proxy object? (not functional yet)
  debug      logical; print debug info?

Details

  it is assumed that the first two dimensions are easting / northing
read_ncdf

Read NetCDF into stars object

Description

Read data from a file (or source) using the NetCDF library directly.

Usage

```
read_ncdf(
  .x, 
  ..., 
  var = NULL, 
  ncsub = NULL, 
  curvilinear = character(0), 
  eps = sqrt(.Machine$double.eps), 
  ignore_bounds = FALSE, 
  make_time = TRUE, 
  make_units = TRUE, 
  proxy = NULL, 
  downsample = 0
)
```

Arguments

- **.x** NetCDF file or source as a character vector or an nc_proxy object.
- **...** ignored
- **var** variable name or names (they must be on matching grids)
- **ncsub** matrix of start, count columns (see Details)
- **curvilinear** length two character named vector with names of variables holding longitude and latitude values for all raster cells. 'stars' attempts to figure out appropriate curvilinear coordinates if they are not supplied.
- **eps** numeric; dimension value increases are considered identical when they differ less than eps
- **ignore_bounds** logical; should bounds values for dimensions, if present, be ignored?
- **make_time** if TRUE (the default), an attempt is made to provide a date-time class from the "time" variable
- **make_units** if TRUE (the default), an attempt is made to set the units property of each variable
- **proxy** logical; if TRUE, an object of class stars_proxy is read which contains array metadata only; if FALSE the full array data is read in memory. If not set, defaults to TRUE when the number of cells to be read is larger than options(stars.n_proxy), or to 1e8 if that option was not set.
**downsample** integer; number of cells to omit between samples along each dimension. e.g. c(1,1,2) would return every other cell in x and y and every third cell in the third dimension (z or t). If 0, no downsampling is applied. Note that this transformation is applied AFTER NetCDF data are read using st_downsample. As such, if proxy=TRUE, this option is ignored.

**Details**

The following logic is applied to coordinates. If any coordinate axes have regularly spaced coordinate variables they are reduced to the offset/delta form with 'affine = c(0, 0)', otherwise the values of the coordinates are stored and used to define a rectilinear grid.

If the data has two or more dimensions and the first two are regular they are nominated as the 'raster' for plotting.

If the `curvilinear` argument is used it specifies the 2D arrays containing coordinate values for the first two dimensions of the data read. It is currently assumed that the coordinates are 2D and that they relate to the first two dimensions in that order.

If `var` is not set the first set of variables on a shared grid is used.

`start` and `count` columns of `ncsub` must correspond to the variable dimension (`nrows`) and be valid index using `var.get.nc` convention (start is 1-based). If the `count` value is `NA` then all steps are included. Axis order must match that of the variable/s being read.

**Examples**

```r
def <- system.file("nc/reduced.nc", package = "stars")
if (require(ncmeta, quietly = TRUE)) {
  read_ncdf(f)
  read_ncdf(f, var = c("anom"))
  read_ncdf(f, nsub = cbind(start = c(1, 1, 1), count = c(10, 12, 1, 1)))
}

if (require(ncmeta, quietly = TRUE)) {
  # precipitation data in a curvilinear NetCDF
  precip_file = system.file("nc/test_stageiv_xyt.nc", package = "stars")
  precip = read_ncdf(precip_file, curvilinear = c("lon", "lat"), ignore_bounds = TRUE)
}

##plot(prec) ## gives error about unique breaks
## remove NAs, zeros, and give a large number
## of breaks (used for validating in detail)
qu_0_omit = function(x, ..., n = 22) {
  x = units::drop_units(na.omit(x))
  c(0, quantile(x[x > 0], seq(0, 1, length.out = n)))
}

if (require(dplyr, quietly = TRUE)) {
  prec_slice = slice(prec, index = 17, along = "time")
  plot(prec_slice, border = NA, breaks = qu_0_omit(prec_slice[[1]]), reset = FALSE)
  nc = sf::read_sf(system.file("gpkg/nc.gpkg", package = "sf"), "nc.gpkg")
  plot(st_geometry(nc), add = TRUE, reset = FALSE, col = NA)
}
Description

read raster/array dataset from file or connection

Usage

read_stars(.x, . . ., options = character(0), driver = character(0), sub = TRUE, quiet = FALSE, NA_value = NA_real_, along = NA_integer_, RasterIO = list(), proxy = is_functions(.x) || (!length(curvilinear) & & is_big(.x, sub = sub, driver = driver, normalize_path = normalize_path, . . .)), curvilinear = character(0), normalize_path = TRUE, RAT = character(0), tolerance = 1e-10, exclude = "")

is_big(x, . . ., sub = sub, n_proxy = options("stars_n_proxy")[[1]] %||% 1e+08)

Arguments

.x character vector with name(s) of file(s) or data source(s) to be read, or a function that returns such a vector

... passed on to st_as_stars if curvilinear was set

options character; opening options

driver character; driver to use for opening file. To override fixing for subdatasets and autodetect them as well, use NULL.

sub character, integer or logical; name, index or indicator of sub-dataset(s) to be read

quiet logical; print progress output?

NA_value numeric value to be used for conversion into NA values; by default this is read from the input file

along length-one character or integer, or list; determines how several arrays are combined, see Details.
list with named parameters for GDAL's RasterIO, to further control the extent, resolution and bands to be read from the data source; see details.

logical; if TRUE, an object of class stars_proxy is read which contains array metadata only; if FALSE the full array data is read in memory. Always FALSE for curvilinear grids. If not set, defaults to TRUE when the number of cells to be read is larger than options(stars.n_proxy), or to 1e8 if that option was not set.

length two character vector with names of subdatasets holding longitude and latitude values for all raster cells, or named length 2 list holding longitude and latitude matrices; the names of this list should correspond to raster dimensions referred to

logical; if FALSE, suppress a call to normalizePath on .x

character; raster attribute table column name to use as factor levels

numeric; passed on to all.equal for comparing dimension parameters.

character; vector with category value(s) to exclude

object to be read with read_stars

integer; number of cells above which .x will be read as stars proxy object, i.e. not as in-memory arrays but left on disk

In case .x contains multiple files, they will all be read and combined with c.stars. Along which dimension, or how should objects be merged? If along is set to NA it will merge arrays as new attributes if all objects have identical dimensions, or else try to merge along time if a dimension called time indicates different time stamps. A single name (or positive value) for along will merge along that dimension, or create a new one if it does not already exist. If the arrays should be arranged along one of more dimensions with values (e.g. time stamps), a named list can passed to along to specify them; see example.

RasterIO is a list with zero or more of the following named arguments: nXOff, nYOff (both 1-based: the first row/col has offset value 1), nXSize, nYSize, nBufXSize, nBufYSize, bands, resample. See https://gdal.org/doxygen/classGDALDataset.html for their meaning; bands is an integer vector containing the band numbers to be read (1-based: first band is 1). Note that if nBufXSize or nBufYSize are specified for downsampling an image, resulting in an adjusted geotransform. resample reflects the resampling method and has to be one of: "nearest_neighbour" (the default), "bilinear", "cubic", "cubic_spline", "lanczos", "average", "mode", or "Gauss".

Data that are read into memory (proxy=FALSE) are read into a numeric (double) array, except for categorical variables which are read into an numeric (integer) array of class factor.

object of class stars

```r
tif = system.file("tif/L7_ETMs.tif", package = "stars")
(x1 = read_stars(tif))
(x2 = read_stars(c(tif, tif)))
```
(x3 = read_stars(c(tif, tif), along = "band"))
(x4 = read_stars(c(tif, tif), along = "new_dimensions")) # create 4-dimensional array
x1o = read_stars(tif, options = "OVERVIEW_LEVEL=1")
t1 = as.Date("2018-07-31")
# along is a named list indicating two dimensions:
read_stars(c(tif, tif, tif, tif), along = list(foo = c("bar1", "bar2"), time = c(t1, t1+2)))

m = matrix(1:120, nrow = 12, ncol = 10)
dim(m) = c(x = 10, y = 12) # named dim
st = st_as_stars(m)
attr(st, "dimensions")$y$delta = -1
attr(st, "dimensions")$y$offset = 12
st
tmp = tempfile(fileext = ".tif")
write_stars(st, tmp)
(red <- read_stars(tmp))
read_stars(tmp, RasterIO = list(nXOff = 1, nYOff = 1, nXSize = 10, nYSize = 12,
nBufXSize = 2, nBufYSIZE = 2))[[1]]
(read <- read_stars(tmp, RasterIO = list(nXOff = 1, nYOff = 1, nXSize = 10, nYSize = 12,
nBufXSize = 2, nBufYSIZE = 2)))
red[[1]] # cell values of subsample grid:
## Not run:
plot(st, reset = FALSE, axes = TRUE, ylim = c(-.1,12.1), xlim = c(-.1,10.1),
main = "nBufXSize & nBufYSize demo", text_values = TRUE)
plot(st_as_sfc(red, as_points = TRUE), add = TRUE, col = "red", pch = 16)
plot(st_as_sfc(st_as_stars(st), as_points = FALSE), add = TRUE, border = 'grey')
plot(st_as_sfc(red, as_points = FALSE), add = TRUE, border = 'green', lwd = 2)
## End(Not run)
file.remove(tmp)

redimension  redimension array, or collapse attributes into a new dimension

Description
redimension array, or collapse attributes into a new dimension

Usage
st_redimension(x, new_dims, along, ...)

## S3 method for class 'stars'
st_redimension(
  x,
  new_dims = st_dimensions(x),
  along = list(new_dim = names(x)),
  ...
)
stars_sentinel2

```r
## S3 method for class 'stars_proxy'
st_redimension(
x,
  new_dims = st_dimensions(x),
  along = list(new_dim = names(x)),
  ...
)
```

### Arguments

- **x**: object of class `stars`
- **new_dims**: target dimensions: either a `dimensions` object or an integer vector with the dimensions’ sizes
- **along**: named list with new dimension name and values
- **...**: ignored

---

**stars_sentinel2**  
*Sentinel-2 sample tile*

---

**Description**

Sentinel-2 sample tile, downloaded from https://scihub.copernicus.eu/ reads the four 10-m bands: B2 (490 nm), B3 (560 nm), B4 (665 nm) and B8 (842 nm)

**Usage**

```r
stars_sentinel2
```

**Format**

An object of class `stars_proxy` (inherits from `stars`) of dimension 10980 x 10980 x 4.

---

**stars_subset**  
*subset stars objects*

---

**Description**

subset stars objects
Usage

```r
## S3 method for class 'stars'
x[i = TRUE, ..., drop = FALSE, crop = !is_curvilinear(x)]

## S3 replacement method for class 'stars'
x[i, downsample = 0] <- value

st_flip(x, which = 1)
```

Arguments

- `x` object of class `stars`
- `i` first selector: integer, logical or character vector indicating attributes to select, or object of class `sf` or `sfc` used as spatial selector; see details
- `...` further (logical or integer vector) selectors, matched by order, to select on individual dimensions
- `drop` logical; if TRUE, degenerate dimensions (with only one value) are dropped
- `crop` logical; if TRUE and parameter `i` is a spatial geometry (`sf` or `sfc`) object, the extent (bounding box) of the result is cropped to match the extent of `i` using `st_crop`. Cropping curvilinear grids is not supported.
- `downsample` downsampling rate used in case `i` is a `stars_proxy` object
- `value` array of dimensions equal to those in `x`, or a vector or value that will be recycled to such an array
- `which` character or integer; dimension(s) to be flipped

Details

if `i` is an object of class `sf`, `sfc` or `bbox`, the spatial subset covering this geometry is selected, possibly followed by cropping the extent. Array values for which the cell centre is not inside the geometry are assigned NA.

in an assignment (or replacement form, `[<-`), argument `i` needs to be a `stars` object with dimensions identical to `x`, and `value` will be recycled to the dimensions of the arrays in `x`.

Value

`st_flip` flips (reverts) the array values along the chosen dimension without(s) changing the dimension properties

Examples

```r
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
x[,,1:3] # select bands
x[,1:100,100:200,] # select x and y by range
x["L7_ETMs.tif"] # select attribute
xy = structure(list(x = c(293253.999046018, 296400.196497684, 296400.196497684), y = c(9113801.64775462, 9111328.49619133)), .Names = c("x", "y"))
```
pts = st_as_sf(data.frame(do.call(cbind, xy)), coords = c("x", "y"), crs = st_crs(x))
image(x, axes = TRUE)
plot(st_as_sfc(st_bbox(pts)), col = NA, add = TRUE)
bb = st_bbox(pts)
(xx = x[bb])
image(xx)
plot(st_as_sfc(bb), add = TRUE, col = NA)
image(x)
pt = st_point(c(x = 290462.103109179, y = 9114202.32594085))
buf = st_buffer(st_sfc(pt, crs = st_crs(x)), 1500)
plot(buf, add = TRUE)

buf = st_sfc(st_polygon(list(st_buffer(pt, 1500)[[1]], st_buffer(pt, 1000)[[1]])),
crs = st_crs(x))
image(x[buf])
plot(buf, add = TRUE, col = NA)
image(x[buf, crop=FALSE])
plot(buf, add = TRUE, col = NA)
lc = read_stars(system.file("tif/lc.tif", package = "stars"))
x = c(orig = lc,
   flip_x = st_flip(lc, "x"),
   flip_y = st_flip(lc, "y"),
   flip_xy = st_flip(lc, c("x", "y")),
   along = 3)
plot(x)

---

**st_apply**

---

**st_apply** apply a function to one or more array dimensions

**Description**

st_apply apply a function to array dimensions: aggregate over space, time, or something else

**Usage**

```r
## S3 method for class 'stars'
st_apply(
  X,
  MARGIN,
  FUN,
  ..., 
  CLUSTER = NULL,
  PROGRESS = FALSE,
  FUTURE = FALSE,
  rename = TRUE,
  .fname,
  single_arg = has_single_arg(FUN, list(...)) || can_single_arg(FUN),
  keep = FALSE
)
```

```r
pts = st_as_sf(data.frame(do.call(cbind, xy)), coords = c("x", "y"), crs = st_crs(x))
image(x, axes = TRUE)
plot(st_as_sfc(st_bbox(pts)), col = NA, add = TRUE)
bb = st_bbox(pts)
(xx = x[bb])
image(xx)
plot(st_as_sfc(bb), add = TRUE, col = NA)
image(x)
pt = st_point(c(x = 290462.103109179, y = 9114202.32594085))
buf = st_buffer(st_sfc(pt, crs = st_crs(x)), 1500)
plot(buf, add = TRUE)

buf = st_sfc(st_polygon(list(st_buffer(pt, 1500)[[1]], st_buffer(pt, 1000)[[1]])),
crs = st_crs(x))
image(x[buf])
plot(buf, add = TRUE, col = NA)
image(x[buf, crop=FALSE])
plot(buf, add = TRUE, col = NA)
lc = read_stars(system.file("tif/lc.tif", package = "stars"))
x = c(orig = lc,
   flip_x = st_flip(lc, "x"),
   flip_y = st_flip(lc, "y"),
   flip_xy = st_flip(lc, c("x", "y")),
   along = 3)
plot(x)
```
Arguments

X object of class stars
MARGIN see apply; index number(s) or name(s) of the dimensions over which FUN will be applied
FUN see apply and see Details.
... arguments passed on to FUN
CLUSTER cluster to use for parallel apply; see makeCluster
PROGRESS logical; if TRUE, use pbapply::pbapply to show progress bar
FUTURE logical; if TRUE, use future.apply::future_apply
rename logical; if TRUE and X has only one attribute and FUN is a simple function name, rename the attribute of the returned object to the function name
.fname function name for the new attribute name (if one or more dimensions are reduced) or the new dimension (if a new dimension is created); if missing, the name of FUN is used
single_arg logical; if TRUE, FUN takes a single argument (like fnNdvi1 below), if FALSE FUN takes multiple arguments (like fnNdvi2 below).
keep logical; if TRUE, preserve dimension metadata (e.g. time stamps)

Details

FUN is a function which either operates on a single object, which will be the data of each iteration step over dimensions MARGIN, or a function that has as many arguments as there are elements in such an object. See the NDVI examples below. The second form can be VERY much faster e.g. when a trivial function is not being called for every pixel, but only once (example).

The heuristics for the default of single_arg work often, but not always; try setting this to the right value when st_apply gives an error.

Value

object of class stars with accordingly reduced number of dimensions; in case FUN returns more than one value, a new dimension is created carrying the name of the function used; see the examples. Following the logic of apply, This new dimension is put before the other dimensions; use aperm to rearrange this, see last example.

Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
st_apply(x, 1:2, mean) # mean band value for each pixel
st_apply(x, c("x", "y"), mean) # equivalent to the above
st_apply(x, 3, mean) # mean of all pixels for each band
## Not run:
st_apply(x, "band", mean) # equivalent to the above
st_apply(x, 1:2, range) # min and max band value for each pixel
fnNdvi1 = function(x) (x[4]-x[3])/(x[4]+x[3]) # ONE argument: will be called for each pixel
fnNdvi2 = function(red,nir) (nir-red)/(nir+red) # n arguments: will be called only once
ndvi1 = st_apply(x, 1:2, fn_ndvi1)
# note that we can select bands 3 and 4 in the first argument:
ndvi2 = st_apply(x[,3:4], 1:2, fn_ndvi2)
all.equal(ndvi1, ndvi2)
# compute the (spatial) variance of each band; https://github.com/r-spatial/stars/issues/430
st_apply(x, 3, function(x) var(as.vector(x))) # as.vector is required!
# to get a progress bar also in non-interactive mode, specify:
if (require(pbapply)) { # install it, if FALSE
  pboptions(type = "timer")
}
st_apply(x, 1:2, range) # dimension "range" is first; rearrange by:
st_apply(x, 1:2, range) %>% aperm(c(2,3,1))

## End(Not run)

---

**st_as_sf**

*Convert stars object into an sf object*

**Description**

Convert stars object into an sf object

**Usage**

```r
## S3 method for class 'stars'
st_as_sfc(x, ..., as_points, which = seq_len(prod(dim(x)[1:2])))
```

```r
## S3 method for class 'stars'
st_as_sf(
  x,
  ..., 
  as_points = FALSE,
  merge = FALSE,
  na.rm = TRUE,
  use_integer = is.logical(x[[1]]) || is.integer(x[[1]]),
  long = FALSE,
  connect8 = FALSE
)
```

```r
## S3 method for class 'stars_proxy'
st_as_sf(x, ..., downsample = 0)
```

**Arguments**

- **x**  
  object of class stars

- **...**  
  ignored

- **as_points**  
  logical; should cells be converted to points or to polygons? See details.
which
merge
na.rm
use_integer
long
connect8
downsample

which
linear index of cells to keep (this argument is not recommended to be used)
merge
logical; if TRUE, cells with identical values are merged (using GDAL_Polygonize or GDAL_FPolygonize); if FALSE, a polygon for each raster cell is returned; see details
na.rm
logical; should missing valued cells be removed, or also be converted to features?
use_integer
(relevant only if merge is TRUE): if TRUE, before polygonizing values are rounded to 32-bits signed integer values (GDALPolygonize), otherwise they are converted to 32-bit floating point values (GDALFPolygonize).
long
logical; if TRUE, return a long table form sf, with geometries and other dimensions recycled
connect8
logical; if TRUE, use 8 connectedness. Otherwise the 4 connectedness algorithm will be applied.
downsample
see st_as_stars

Details

If merge is TRUE, only the first attribute is converted into an sf object. If na.rm is FALSE, areas with NA values are also written out as polygons. Note that the resulting polygons are typically invalid, and use st_make_valid to create valid polygons out of them.

Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
x = x[,1:100,1:100,6] # subset of a band with lower values in it
x[[1]][x[[1]] < 30] = NA # set lower values to NA
x[[1]] = x[[1]] < 100 # make the rest binary
x
(p = st_as_sf(x)) # removes NA areas
(p = st_as_sf(x[,,,1], merge = TRUE)) # glues polygons together
all(st_is_valid(p)) # not all valid, see details
plot(p, axes = TRUE)
(p = st_as_sf(x, na.rm = FALSE, merge = TRUE)) # includes polygons with NA values
plot(p, axes = TRUE)

st_as_stars

convert objects into a stars object

Description

convert objects into a stars object
Usage

```r
st_as_stars(.x, ...)
```

## S3 method for class 'list'
```r
st_as_stars(.x, ..., dimensions = NULL)
```

## Default S3 method:
```r
st_as_stars(.x = NULL, ..., raster = NULL)
```

## S3 method for class 'stars'
```r
st_as_stars(.x, ..., curvilinear = NULL, crs = st_crs(4326))
```

## S3 method for class 'bbox'
```r
st_as_stars(.x, ..., nx, ny, dx = dy, dy = dx, xlim = .x[c("xmin", "xmax")], ylim = .x[c("ymin", "ymax")], values = 0, n = 64800, pretty = FALSE, inside = FALSE, nz, proxy = FALSE)
```

## S3 method for class 'sf'
```r
st_as_stars(.x, ..., name = attr(.x, "sf_column"))
```

## S3 method for class 'Raster'
```r
st_as_stars(.x, ..., att = 1, ignore_file = FALSE)
```

## S3 method for class 'SpatRaster'
```r
st_as_stars(.x, ..., ignore_file = FALSE, as_attributes = all(terra::is.factor(.x))
```

## S3 method for class 'ncdfgeom'
```r
st_as_stars(.x, ..., sf_geometry = NA)
```

## S3 method for class 'stars_proxy'
```r
```
\begin{verbatim}
st_as_stars(
  .x, 
  ..., 
  downsample = 0, 
  url = attr(.x, "url"), 
  envir = parent.frame()
)

## S3 method for class 'data.frame'
st_as_stars(
  .x, 
  ..., 
  dims = coords, 
  xy = dims[1:2], 
  y_decreasing = TRUE, 
  coords = 1:2
)

## S3 method for class 'xts'
st_as_stars(.x, ..., dimensions, name = "attr")

## S3 method for class 'OpenStreetMap'
st_as_stars(.x, ..., as_col = FALSE)

Arguments

.x object to convert
...
in case .x is of class bbox, arguments passed on to pretty. In case .x is of class nc_proxy, arguments passed on to \texttt{read_ncdf}.
dimensions object of class dimensions
raster character; the names of the dimensions that denote raster dimensions
curvilinear only for creating curvilinear grids: named length 2 list holding longitude and latitude matrices; the names of this list should correspond to raster dimensions referred to
crs object of class \texttt{crs} with the coordinate reference system of the values in \texttt{curvilinear}; see details
nx integer; number of cells in x direction; see details
ny integer; number of cells in y direction; see details
dx numeric or object of class units; cell size in x direction; see details
dy numeric or object of class units; cell size in y direction; see details
xlim length 2 numeric vector with extent (min, max) in x direction
ylim length 2 numeric vector with extent (min, max) in y direction
values value(s) to populate the raster values with
n the (approximate) target number of grid cells
pretty logical; should cell coordinates have pretty values?
\end{verbatim}
inside logical; should all cells entirely fall inside the bbox, potentially not covering it completely (TRUE), or always cover the bbox (FALSE), or find a good approximation (NA, default)?

nz integer; number of cells in z direction; if missing no z-dimension is created.

proxy logical; should a stars_proxy object be created? (requires gdal_create binary when sf < 1.0-6)

name character; attribute name for array from an xts object

att see factorValues; column in the RasterLayer’s attribute table

ignore_file logical; if TRUE, ignore the SpatRaster object file name

as_attributes logical; if TRUE and .x has more than one layer, load these as separate attributes rather than as a band or time dimension (only implemented for the case where ignore_file is TRUE)

sf_geometry sf data.frame with geometry and attributes to be added to stars object. Must have same number of rows as timeseries instances.

downsampling integer: if larger than 0, downsample with this rate (number of pixels to skip in every row/column); if length 2, specifies downsampling rate in x and y.

url character; URL of the stars endpoint where the data reside

envir environment to resolve objects in

dims the column names or indices that form the cube dimensions

xy the x and y raster dimension names or indices; only takes effect after dims has been specified

y_decreasing logical; if TRUE, (numeric) y values get a negative delta (decrease with increasing index)

cords same as dims, for symmetry with st_as_sf

as_col logical; return rgb numbers (FALSE) or (character) color values (TRUE)?

Details

if curvilinear is a stars object with longitude and latitude values, its coordinate reference system is typically not that of the latitude and longitude values.

For the bbox method: if pretty is TRUE, raster cells may extend the coordinate range of .x on all sides. If in addition to nx and ny, dx and dy are also missing, these are set to a single value computed as \(\sqrt{\text{diff}(\text{xlim})\times\text{diff}(\text{ylim})/n}\). If nx and ny are missing, they are computed as the (ceiling, floor, or rounded to integer value) of the ratio of the (x or y) range divided by (dx or dy), depending on the value of inside. Positive dy will be made negative. Further named arguments (...) are passed on to pretty. If dx or dy are units objects, their value is converted to the units of st_crs(.x) (only when sf >= 1.0-7).

For the ncdfgeom method: objects are point-timeseries with optional line or polygon geometry for each timeseries specified with the sf_geometry parameter. See ncdfgeom for more about this NetCDF-based format for geometry and timeseries.

for the xts methods, if dimensions are provided, time has to be the first dimension.
Examples

```r
if (require(plm, quietly = TRUE)) {
  data(Produc, package = "plm")
  st_as_stars(Produc, y_decreasing = FALSE)
  data(Produc, package = "plm")
  st_as_stars(Produc, y_decreasing = FALSE)
}
```

---

**st_contour**

Compute or plot contour lines or sets

**Description**

Compute contour lines or sets

**Usage**

```r
st_contour(
  x,
  na.rm = TRUE,
  contour_lines = FALSE,
  breaks = classInt::classIntervals(na.omit(as.vector(x[[1]])))$brks
)
```

**Arguments**

- `x` object of class `stars`
- `na.rm` logical; should missing valued cells be removed, or also be converted to features?
- `contour_lines` logical; if FALSE, polygons are returned (contour sets), otherwise contour lines
- `breaks` numerical; values at which to "draw" contour levels

**Details**

this function requires GDAL >= 2.4.0

**See Also**

for polygonizing rasters following grid boundaries, see `st_as_sf` with arguments `as_points=FALSE` and `merge=TRUE`; `contour` plots contour lines using R’s native algorithm (which also plots contour levels)
st_coordinates

---

**Description**

retrieve coordinates for raster or vector cube cells

**Usage**

```r
## S3 method for class 'stars'
st_coordinates(x, ..., add_max = FALSE, center = TRUE)

## S3 method for class 'stars'
as.data.frame(x, ..., add_max = FALSE, center = NA)

as_tibble.stars(.x, ..., add_max = FALSE, center = NA)
```

**Arguments**

- `x` object of class stars
- `...` ignored
- `add_max` logical; if TRUE, dimensions are given with a min (x) and max (x_max) value
- `center` logical; (only if add_max is FALSE): should grid cell center coordinates be returned (TRUE) or offset values (FALSE)? center can be a named logical vector or list to specify values for each dimension.
- `.x` object to be converted to a tibble

---

st_crop

---

**Description**

crop a stars object

**Usage**

```r
## S3 method for class 'stars_proxy'
st_crop(
  x,
  y,
  ...,  
crop = TRUE,
epsilon = sqrt(.Machine$double.eps),
collect = TRUE
```

---
## S3 method for class 'stars'
st_crop(
  x,
  y,
  ..., 
  crop = TRUE,
  epsilon = sqrt(.Machine$double.eps),
  as_points = all(st_dimension(y) == 2, na.rm = TRUE)
)

### Arguments

- **x**: object of class stars
- **y**: object of class sf, sfc or bbox; see Details below.
- **...**: ignored
- **crop**: logical; if TRUE, the spatial extent of the returned object is cropped to still cover obj, if FALSE, the extent remains the same but cells outside y are given NA values.
- **epsilon**: numeric; factor to shrink the bounding box of y towards its center before cropping.
- **collect**: logical; if TRUE, repeat cropping on stars object, i.e. after data has been read
- **as_points**: logical; only relevant if y is of class sf or sfc: if FALSE, treat x as a set of points, else as a set of small polygons. Default: TRUE if y is two-dimensional, else FALSE; see Details

### Details

For raster x, st_crop selects cells that intersect with y. For intersection, are raster cells interpreted as points or as small polygons? If y is of class stars, x raster cells are interpreted as points; if y is of class bbox, x cells are interpreted as cells (small polygons). Otherwise, if as_points is not given, cells are interpreted as points if y has a two-dimensional geometry.

### Examples

```r
l7 = read_stars(system.file("tif/L7_ETMs.tif", package = "stars"))
d = st_dimensions(l7)

# area around cells 3:10 (x) and 4:11 (y):  
offset = c(d["x"]$offset, d["y"]$offset)
res = c(d["x"]$delta, d["y"]$delta)
bb = st_bbox(c(xmin = offset[1] + 2 * res[1],  
ymin = offset[2] + 11 * res[2],  
xmax = offset[1] + 10 * res[1],  
ymax = offset[2] + 3 * res[2]), crs = st_crs(l7))

l7[bb]
# equivalent:
st_crop(l7, bb)
```
st_dimensions

```
plot(l7[,1:13,1:13,1], reset = FALSE)
image(l7[bb,,,1], add = TRUE, col = sf.colors())
plot(st_as_sfc(bb), add = TRUE, border = 'green', lwd = 2)

# slightly smaller bbox:
bb = st_bbox(c(xmin = offset[1] + 2.1 * res[1],
ymin = offset[2] + 10.9 * res[2],
xmax = offset[1] + 9.9 * res[1],
l7[bb]

plot(l7[,1:13,1:13,1], reset = FALSE)
image(l7[bb,,,1], add = TRUE, col = sf.colors())
plot(st_as_sfc(bb), add = TRUE, border = 'green', lwd = 2)

# slightly larger bbox:
bb = st_bbox(c(xmin = offset[1] + 1.9 * res[1],
xmax = offset[1] + 10.1 * res[1],
ymax = offset[2] + 2.9 * res[2]), crs = st_crs(l7))
l7[bb]

plot(l7[,1:13,1:13,1], reset = FALSE)
image(l7[bb,,,1], add = TRUE, col = sf.colors())
plot(st_as_sfc(bb), add = TRUE, border = 'green', lwd = 2)

# half a cell size larger bbox:
bb = st_bbox(c(xmin = offset[1] + 1.49 * res[1],
ymin = offset[2] + 11.51 * res[2],
xmax = offset[1] + 10.51 * res[1],
ymax = offset[2] + 2.49 * res[2]), crs = st_crs(l7))
l7[bb]

plot(l7[,1:13,1:13,1], reset = FALSE)
image(l7[bb,,,1], add = TRUE, col = sf.colors())
plot(st_as_sfc(bb), add = TRUE, border = 'green', lwd = 2)
```

---

**Description**

get dimensions from stars object

**Usage**

```
st_dimensions(.x, ...)  
```

```
## S3 method for class 'stars'
```
st_dimensions(.x, ...)

st_dimensions(x) <- value

## S3 replacement method for class 'stars'
st_dimensions(x) <- value

## S3 replacement method for class 'stars_proxy'
st_dimensions(x) <- value

## S3 replacement method for class 'list'
st_dimensions(x) <- value

## S3 method for class 'array'
st_dimensions(.x, ...)

## Default S3 method:
st_dimensions(
  .x,
  ..., .raster,
  affine = c(0, 0),
  cell_midpoints = FALSE,
  point = FALSE
)

st_set_dimensions(
  .x,
  which,
  values = NULL,
  point = NULL,
  names = NULL,
  xy,
  ...
)

st_get_dimension_values(.x, which, ..., where = NA, max = FALSE, center = NA)

Arguments

.x object to retrieve dimensions information from

... further arguments

x object of class dimensions

value new object of class dimensions, with matching dimensions

.raster length 2 character array with names (if any) of the raster dimensions

affine numeric; specify parameters of the affine transformation
cell_midpoints logical; if TRUE AND the dimension values are strictly regular, the values are interpreted as the cell midpoint values rather than the cell offset values when calculating offset (i.e., the half-cell-size correction is applied); can have a value for each dimension, or else is recycled

point logical; does the pixel value (measure) refer to a point (location) value or to an pixel (area) summary value?

which integer or character; index or name of the dimension to be changed

values values for this dimension (e.g. sfc list-column), or length-1 dimensions object

names character; vector with new names for all dimensions, or with the single new name for the dimension indicated by which

xy length-2 character vector; (new) names for the x and y raster dimensions

where character, one of 'start', 'center' or 'end'. Set to NA (default) to ignore and use max and center explicitly. This argument provides a convenient alternative to setting max and center.

max logical; if TRUE return the end, rather than the beginning of an interval

center logical; if TRUE return the center of an interval; if NA return the center for raster dimensions, and the start of intervals in other cases

details dimensions can be specified in two ways. The simplest is to pass a vector with numeric values for a numeric dimension, or character values for a categorical dimension. Parameter cell_midpoints is used to specify whether numeric values refer to the offset (start) of a dimension interval (default), or to the center; the center case is only available for regular dimensions. For rectilinear numeric dimensions, one can specify either a vector with cell borders (start values), or a data.frame with two columns named "start" and "end", with the respective interval start and end values. In the first case, the end values are computed from the start values by assuming the last two intervals have equal width.

value the dimensions attribute of x, of class dimensions

examples

```r
x = read_stars(system.file("tif/L7_ETMs.tif", package = "stars"))
# Landsat 7 ETM+ band semantics: https://landsat.gsfc.nasa.gov/the-enhanced-thematic-mapper-plus/
# set bands to values 1,2,3,4,5,7:
(x1 = st_set_dimensions(x, "band", values = c(1,2,3,4,5,7), names = "band_number", point = TRUE))
# set set bands as bandwidth
(rbind(c(0.45,0.515), c(0.525,0.605), c(0.63,0.69), c(0.775,0.90), c(1.55,1.75), c(2.08,2.35))) %>%
  units::set_units("um") -> bw # or: units::set_units(bw[,1]) + bw[,2]
# set bandwidth midpoint:
(x2 = st_set_dimensions(x, "band", values = 0.5 * (bw[,1]+bw[,2]),
  names = "bandwidth_midpoint", point = TRUE))
# set bandwidth intervals:
(x3 = st_set_dimensions(x, "band", values = make_intervals(bw), names = "bandwidth")
```

m = matrix(1:20, nrow = 5, ncol = 4)
\begin{verbatim}
dim(m) = c(x = 5, y = 4) # named dim (s = st_as_stars(m))
st_get_dimension_values(s, 'x', where = "start")
st_get_dimension_values(s, 'x', center = FALSE)
st_get_dimension_values(s, 'x', where = "center")
st_get_dimension_values(s, 'x', center = TRUE)
st_get_dimension_values(s, 'x', where = "end")
st_get_dimension_values(s, 'x', max = TRUE)
\end{verbatim}

---

**st_dim_to_attr**

create an array with dimension values

**Description**

create an array with dimension values

**Usage**

\[
st\_dim\_to\_attr(x, \text{which} = \text{seq\_along(dim(x))})
\]

**Arguments**

- **x**: object of class `stars`
- **which**: integer; indices of the dimensions to address (default: all)

**Value**

`stars` object with dimension values as attributes

**Examples**

\[
tif = \text{system.file("tif/L7\_ETMs.tif", package = "stars")}
x1 = \text{read\_stars(tif)}
(x = st\_dim\_to\_attr(x1))
plot(x)
(x = st\_dim\_to\_attr(x1, 2:3))
plot(x)
(x = st\_dim\_to\_attr(x1, 3))
plot(x)
\]
**st_downsample**

downsample stars or stars_proxy object by skipping rows, columns and bands

**Description**

downsample stars or stars_proxy object by skipping rows, columns and bands

**Usage**

```r
st_downsample(x, n, ...)
```

## S3 method for class 'stars'

```r
st_downsample(x, n, ...)
```

## S3 method for class 'stars_proxy'

```r
st_downsample(x, n, ...)
```

**Arguments**

- `x` object of class stars or stars_proxy
- `n` numeric; the number of pixels/lines/bands etc that will be skipped; see Details.
- `...` ignored

**Details**

If all n == 0, no downsampling takes place; if it is 1, every second row/column/band is skipped, if it is 2, every second+third row/column/band are skipped, etc.

Downsampling a stars_proxy object returns a stars object, is equivalent to calling `st_as_stars(x, downsample = 2)`, and only downsamples the first two (x and y) dimensions.

Downsampled regular rasters keep their dimension offsets, have a cell size (delta) that is n[i]+1 times larger, and may result in a (slightly) different extent.

**st_extract**

Extract cell values at point locations

**Description**

Extract cell values at point locations
Usage

```r
st_extract(x, ...)
```

## S3 method for class 'stars'

```r
st_extract(
  x,
  at,
  ..., 
  bilinear = FALSE,
  time_column = attr(at, "time_column") %||% attr(at, "time_col"),
  interpolate_time = bilinear,
  FUN = mean
)
```

### Arguments

- **x**: object of class `stars` or `stars_proxy`
- **...**: passed on to `aggregate.stars` when geometries are not exclusively POINT geometries
- **at**: object of class `sf` or `sfc` with geometries, or two-column matrix with points in rows, indicating where to extract `x`
- **bilinear**: logical; use bilinear interpolation rather than nearest neighbour?
- **time_column**: character or integer; name or index of a column with time or date values that will be matched to values of the dimension "time" in `x`, after which this dimension is reduced. This is useful to extract data cube values along a trajectory; see [https://github.com/r-spatial/stars/issues/352](https://github.com/r-spatial/stars/issues/352).
- **interpolate_time**: logical; should time be interpolated? if FALSE, time instances are matched using the coinciding or the last preceding time in the data cube.
- **FUN**: function used to aggregate pixel values when geometries of `at` intersect with more than one pixel

### Details

Points outside the raster are returned as NA values. For large sets of points for which extraction is needed, passing a matrix as to `at` may be much faster than passing an `sf` or `sfc` object.

### Value

If `at` is of class `matrix`, a matrix with extracted values is returned; otherwise: if `x` has more dimensions than only `x` and `y` (raster), an object of class `stars` with POINT geometries replacing `x` and `y` raster dimensions, if this is not the case, an object of `sf` with extracted values.

### Examples

```r
tif = system.file("tif/L7_ETMs.tif", package = "stars")
r = read_stars(tif)
```
st_intersects.stars

```r
pnt = st_sample(st_as_sfc(st_bbox(r)), 10)
st_extract(r, pnt)
st_extract(r, pnt) %>% st_as_sf()
st_extract(r[, , , , 1], pnt)
st_extract(r, st_coordinates(pnt)) # "at" is a matrix: return a matrix
```

---

**st_intersects.stars**  
*spatial intersect predicate for stars and sfc object*

**Description**

spatial intersect predicate for stars and sfc object

**Usage**

```r
## S3 method for class 'stars'
st_intersects(x, y, sparse = TRUE, ..., as_points = NA, transpose = FALSE)
```

**Arguments**

- `x`  
  object of class stars

- `y`  
  object that has an ‘st_geometry’ method: of class ‘sf’ or ‘sfc’, or ‘stars’ object with an ‘sfc’ dimension

- `sparse`  
  logical; if TRUE, return the a sparse logical matrix (object of class ‘sgbp’), if FALSE, return a logical matrix

- `...`  
  ignored, or passed on to ‘st_intersects.sf’ for curvilinear grids

- `as_points`  
  logical, should grid cells be considered as points (TRUE) or polygons (FALSE)? Default: FALSE and warning emitted

- `transpose`  
  logical; should the transpose of the ‘sgbp’ object be returned?

**Details**

curvilinear grids are always converted to polygons, so points on grid boundaries may intersect with two cells touched; for other grids each cell boundary or corner belongs only to one cell.

**Value**

‘sgbp’ object if sparse = TRUE, logical matrix otherwise
Spatially join a stars and an 'sf' object

Arguments

- **x**: object of class stars
- **y**: object of class sf, or one that can be coerced into that by \texttt{st\_as\_sf}
- **join**: the join function, which should return an sgbp object; see details
- **what**: "left1", "right" or "inner"; see details
- **as\_points**: logical; controls whether grid cells in \texttt{x} will be treated as points, or as cell areas; the \texttt{st\_intersects.stars} method by default will derive this from \texttt{x}'s metadata, or else assume areas.
- **warn**: logical; if TRUE, warn on 1-to-many matches when \texttt{what} is "left1"

Details

When there is more than one match to a single \texttt{x} value, the first matching record from \texttt{y} is taken (and if \texttt{warn} is TRUE a warning is raised). If what is "inner", an object of class \texttt{sf} with all matching records of \texttt{x} and \texttt{y}.

Value

If what is "left1", an object of class stars with the (first) value of \texttt{y} at spatial instances of \texttt{x}
**Description**

build mosaic (composite) of several spatially disjoint stars objects

**Usage**

```r
st_mosaic(.x, ...)```

### S3 method for class 'stars'

```r
st_mosaic(
  .x,
  ..., 
  dst = tempfile(fileext = file_ext),
  options = c("-vrtnodata", "-9999", "-srcnodata", "nan"),
  file_ext = ".tif"
)
```

### S3 method for class 'character'

```r
st_mosaic(
  .x,
  ..., 
  dst = tempfile(fileext = file_ext),
  options = c("-vrtnodata", "-9999"),
  file_ext = ".tif"
)
```

### S3 method for class 'stars_proxy'

```r
st_mosaic(
  .x,
  ..., 
  dst = tempfile(fileext = file_ext),
  options = c("-vrtnodata", "-9999"),
  file_ext = ".tif"
)
```

**Arguments**

- `.x` object of class stars, or character vector with input dataset names
- `...` further input stars objects
- `dst` character; destination file name
- `options` character; options to the gdalbuildvrt command
- `file_ext` character; file extension, determining the format used to write to (".tif" implies GeoTIFF)
st_rasterize

Details

the `gdal` function `buildvrt` builds a mosaic of input images; these input images can be multi-band, but not higher-dimensional data cubes or stars objects with multiple attributes

uses `gdal_utils` to internally call `buildvrt`; no executables external to R are called.

Value

the stars method returns a stars object with the composite of the input; the character method returns the file name of the file with the mosaic; see also the GDAL documentation of `gdalbuildvrt`

Examples

```r
x = read_stars(system.file("tif/L7_ETMs.tif", package = "stars"))
x1 = x[,100:200,100:200,]
x2 = x[,150:300,150:300,]
plot(st_mosaic(x1, x2))
```

Description

rasterize simple feature geometries

Usage

```r
st_rasterize(
sf,
template = guess_raster(sf, ...) %||% st_as_stars(st_bbox(sf), values = NA_real_, ...),
file = tempfile(),
driver = "GTiff",
options = character(0),
align = FALSE,
proxy = FALSE,
...)
```

Arguments

- `sf` object of class `sf`
- `template` stars object with desired target geometry, or target geometry alignment if `align=TRUE`
- `file` temporary file name
- `driver` driver for temporary file
- `options` character; options vector for GDALRasterize
align logical; if TRUE, template contain the geometry alignment, informing target resolution and offset only.
proxy logical; should a proxy object be returned?
...
arguments passed on to st_as_stars

Examples

demo(nc, echo = FALSE, ask = FALSE)
(x = st_rasterize(nc)) # default grid:
plot(x, axes = TRUE)
# a bit more customized grid:
(x = st_rasterize(nc, st_as_stars(st_bbox(nc), nx = 100, ny = 50, values = NA_real_)))
plot(x, axes = TRUE)
(ls = st_sf(a = 1:2, st_sfc(st_linestring(rbind(c(0.1, 0), c(1.1, 1))),
   st_linestring(rbind(c(0, 0.05), c(1, 0.05))))))
(grd = st_as_stars(st_bbox(ls), nx = 10, ny = 10, xlim = c(0, 1.0), ylim = c(0, 1),
   values = NA_real_))
# Only the left-top corner is part of the grid cell:
sf_extSoftVersion()$"GDAL"
plot(st_rasterize(ls, grd), axes = TRUE, reset = FALSE) # ALL_TOUCHED=FALSE;
plot(ls, add = TRUE, col = "red")
plot(st_rasterize(ls, grd, options = "ALL_TOUCHED=TRUE"), axes = TRUE, reset = FALSE)
plot(ls, add = TRUE, col = "red")
# add lines to existing 0 values, summing values in case of multiple lines:
(grd = st_as_stars(st_bbox(ls), nx = 10, ny = 10, xlim = c(0, 1.0), ylim = c(0, 1), values = 0))
r = st_rasterize(ls, grd, options = c("MERGE_ALG=ADD", "ALL_TOUCHED=TRUE"))
plot(r, axes = TRUE, reset = FALSE)
plot(ls, add = TRUE, reset = FALSE)
plot(r, axes = TRUE, col = "red")

---

st_raster_type get the raster type (if any) of a stars object

Description
get the raster type (if any) of a stars object

Usage

st_raster_type(x, dimension = character(0))

Arguments

x object of class stars
dimension optional: numbers or names of dimension(s) to get per-dimension type

Details
categories "curvilinear" and "affine" only refer to the relationship between a pair of spatial (raster) dimensions.
Value

if dimension is not specified, return the spatial raster type: one of NA (if the object does not have raster dimensions), "curvilinear", "rectilinear", "affine", or "regular". In case dimension(s) are specified, return one of "regular", "rectilinear" (irregular but numeric), or "discrete" (anything else).

Examples

tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
st_raster_type(x)
st_raster_type(x, 1:3)

---

st_rgb  reduce dimension to rgb (alpha) hex values

Description

reduce dimension to rgb (alpha) hex values

Usage

st_rgb(
  x,
  dimension = 3,
  use_alpha = dim(x)[dimension] == 4,
  maxColorValue = 255L,
  probs = c(0, 1),
  stretch = NULL
)

Arguments

x  object of class stars
dimension  dimension name or number to reduce
use_alpha  logical; if TRUE, the fourth band will be used as alpha values
maxColorValue  integer; maximum value for colors
probs  probability values for quantiles used for stretching by "percent".
stretch  logical or character; if TRUE or "percent", each band is stretched to 0 ... maxColorValue by "percent clip" method using probs values. If "histogram", a "histogram equalization" is performed (probs values are ignored). If stretch is NULL or FALSE, no stretching is performed. Other character values are interpreted as "percent" and a message will be printed.
**st_set_bbox**

*set bounding box parameters of regular grid*

### Description

set bounding box parameters of regular grid

### Usage

```r
st_set_bbox(x, value, ...)
```

### Arguments

- **x**: object of class dimensions, stars or stars_proxy
- **value**: object of class bbox
- **...**: ignored

### Details

the dimension’s bands are mapped to red, green, blue, alpha; if a different ordering is wanted, use [.stars to reorder a dimension, see examples

### See Also

- `st_apply`, `rgb`

### Examples

```r
tif = system.file("tif/L7_ETMs.tif", package = "stars")
x = read_stars(tif)
st_rgb(x[,,3:1])
r = st_rgb(x[,,,c(6,5,4,3)], 3, use_alpha=TRUE)  # now R=6,G=5,B=4,alpha=3
if (require(ggplot2)) {
  ggplot() + geom_stars(data = r) + scale_fill_identity()
}
r = st_rgb(x[,,,3:1],
  probs = c(0.01, 0.99),
  stretch = "percent")
plot(r)
r = st_rgb(x[,,,3:1],
  probs = c(0.01, 0.99),
  stretch = "histogram")
plot(r)
```
**st_sfc2xy**

*replace POINT simple feature geometry list with an x y raster*

**Description**

replace POINT simple feature geometry list with an x y raster

**Usage**

```r
st_sfc2xy(x, ...)
```

**Arguments**

- `x` object of class `stars`, or of class `sf`
- `...` passed on to `as.data.frame.stars`

**Value**

object of class `stars` with a POINT list replaced by x and y raster dimensions. This only works when the points are distributed over a regular or rectilinear grid.

---

**st_tile**

*Specify parameters to load raster in blocks*

**Description**

Helper function for specifying the block parameters (nXOff, nYOff, nXsize, and nYSize) required by RasterIO argument in `read_stars`

**Usage**

```r
st_tile(img_rows, img_cols, x_window, y_window, overlap = 0)
```

**Arguments**

- `img_rows` number of input raster rows (integer)
- `img_cols` number of input raster columns (integer)
- `x_window` number of rows in block (integer)
- `y_window` number of columns in block (integer)
- `overlap` number of overlapping pixels (integer)

**Value**

matrix with specified nXOff, nYOff, nXsize, and nYSize parameters for every block
## Examples

```r
# Not run:
tif = system.file("tif/L7_ETMs.tif", package = "stars")
r = read_stars(tif, proxy = TRUE)
tiles = st_tile(nrow(r), ncol(r), 256, 256)
for (i in seq_len(nrow(tiles))) {
  tile = read_stars(tif, proxy = FALSE, RasterIO = tiles[i, ])
  write_tiles to separate files
  write_stars(tile, dsn = paste0(i, ".tif"))
}

# End(Not run)
```

---

### st_transform

**transform geometries in stars objects to a new coordinate reference system, without warping**

### Description

transform geometries in stars objects to a new coordinate reference system, without warping

### Usage

```r
# S3 method for class 'stars'
st_transform(x, crs, ...)
```

```r
# S3 method for class 'stars'
st_transform_proj(x, crs, ...)
```

### Arguments

- **x**: object of class `stars`, with either raster or simple feature geometries
- **crs**: object of class `crs` with target `crs`
- **...**: ignored

### Details

For simple feature dimensions, `st_transform` is called, leading to lossless transformation. For grid-ded spatial data, a curvilinear grid with transformed grid cell (centers) is returned, which is also lossless. To convert this to a regular grid in the new `CRS`, use `st_warp` (which is in general lossy).

### See Also

- `st_warp`
Examples

```r
geomatrix = system.file("tif/geomatrix.tif", package = "stars")
(x = read_stars(geomatrix))
new = st_crs(4326)
y = st_transform(x, new)
plot(st_transform(st_as_sfc(st_bbox(x)), new), col = NA, border = 'red')
plot(st_as_sfc(y, as_points=FALSE), col = NA, border = 'green', axes = TRUE, add = TRUE)
image(y, col = heat.colors(12), add = TRUE)
plot(st_as_sfc(y, as_points=TRUE), pch=3, cex=.5, col = 'blue', add = TRUE)
plot(st_transform(st_as_sfc(x, as_points=FALSE), new), add = TRUE)
```

---

**st_warp**

Warp (resample) grids in stars objects to a new grid, possibly in an new coordinate reference system

**Description**

Warp (resample) grids in stars objects to a new grid, possibly in an new coordinate reference system

**Usage**

```r
st_warp(
  src,
  dest,
  ...,  # ignored
  crs = NA_crs_,
  cellsize = NA_real_,
  segments = 100,
  use_gdal = FALSE,
  options = character(0),
  no_data_value = NA_real_,
  debug = FALSE,
  method = "near",
  threshold = ifelse(is.na(cellsize), Inf, cellsize/2)
)
```

**Arguments**

- `src` : object of class stars with source raster
- `dest` : object of class stars with target raster geometry
- `...` : ignored
- `crs` : coordinate reference system for destination grid, only used when `dest` is missing
- `cellsize` : length 1 or 2 numeric; cellsize in target coordinate reference system units
- `segments` : (total) number of segments for segmentizing the bounding box before transforming to the new crs
- `use_gdal` : logical; if TRUE, use gdal’s warp or warper, through `gdal_utils`
st_warp

options character vector with options, passed on to gdalwarp
no_data_value value used by gdalwarp for no_data (NA) when writing to temporary file; not setting this when use_gdal is TRUE leads to a warning
debug logical; if TRUE, do not remove the temporary gdalwarp destination file, and print its name
method character; see details for options; methods other than near only work when use_gdal=TRUE
threshold numeric; distance threshold for warping curvilinear grids: new cells at distances larger than threshold are assigned NA values.

Details

method should be one of near, bilinear, cubic, cubicspline, lanczos, average, mode, max, min, med, q1 or q3; see https://github.com/r-spatial/stars/issues/109

For gridded spatial data (dimensions x and y), see figure; the existing grid is transformed into a regular grid defined by dest, possibly in a new coordinate reference system. If dest is not specified, but crs is, the procedure used to choose a target grid is similar to that of projectRaster. This entails: (i) the envelope (bounding box polygon) is transformed into the new crs, possibly after segmentation (red box); (ii) a grid is formed in this new crs, touching the transformed envelope on its East and North side, with (if cellsize is not given) a cellsize similar to the cell size of src, with an extent that at least covers x; (iii) for each cell center of this new grid, the matching grid cell of x is used; if there is no match, an NA value is used.

Examples

geomatrix = system.file("tif/geomatrix.tif", package = "stars")
(x = read_stars(geomatrix))
new_crs = st_crs(4326)
y = st_warp(x, crs = new_crs)
plot(st_transform(st_as_sfc(st_bbox(x)), new_crs), col = NA, border = 'red')
plot(st_as_sfc(y, as_points=FALSE), col = NA, border = 'green', axes = TRUE, add = TRUE)
image(y, add = TRUE, nbreaks = 6)
plot(st_as_sfc(y, as_points=TRUE), pch=3, cex=.5, col = 'blue', add = TRUE)
plot(st_transform(st_as_sfc(x, as_points=FALSE), new_crs), add = TRUE)
# warp 0-360 raster to -180-180 raster:
r = read_stars(system.file("nc/reduced.nc", package = "stars"))
r %>% st_set_crs(4326) %>% st_warp(st_as_stars(st_bbox(), dx = 2)) -> s
plot(r, axes = TRUE) # no CRS set, so no degree symbols in labels
plot(s, axes = TRUE)
# downsample raster (90 to 270 m)
r = read_stars(system.file("tif/olinda_dem_utm25s.tif", package = "stars"))
r270 = st_as_stars(st_bbox(r), dx = 270)
r270 = st_warp(r, r270)
st_xy2sfc  
replace x y raster dimensions with simple feature geometry list (points, or polygons = rasterize)

Description  
replace x y raster dimensions with simple feature geometry list (points, or polygons = rasterize)

Usage  
st_xy2sfc(x, as_points, ..., na.rm = TRUE)

Arguments  
x  
object of class stars

as_points  
logical; if TRUE, generate points at cell centers, else generate polygons

...  
arguments passed on to st_as_sfc

na.rm  
logical; omit (remove) cells which are entirely missing valued (across other dimensions)?

Value  
object of class stars with x and y raster dimensions replaced by a single sfc geometry list column containing either points, or polygons. Adjacent cells with identical values are not merged; see st_rasterize for this.

write_stars  
write stars object to gdal dataset (typically: to file)

Description  
write stars object to gdal dataset (typically: to file)

Usage  
write_stars(obj, dsn, layer, ...)

## S3 method for class 'stars'
write_stars(
  obj,
  dsn,
  layer = 1,
  ...,
  driver = detect.driver(dsn),
  options = character(0),
)
write_stars

```
type = if (is.factor(obj[[1]]) && length(levels(obj[[1]])) < 256) "Byte" else "Float32",
      NA_value = NA_real_,
      update = FALSE,
      normalize_path = TRUE
    )

## S3 method for class 'stars_proxy'
write_stars(
  obj,
  dsn,
  layer = 1,
  ...,
  driver = detect.driver(dsn),
  options = character(0),
  type = "Float32",
  NA_value = NA_real_,
  chunk_size = c(dim(obj)[1], floor(2.5e+07/dim(obj)[1])),
  progress = TRUE
)

detect.driver(filename)
```

Arguments

- `obj` object of class `stars`
- `dsn` gdal dataset (file) name
- `layer` attribute name; if missing, the first attribute is written
- `...` passed on to `gdal_write`
- `driver` driver driver name; see `st_drivers`
- `options` character vector with dataset creation options, passed on to GDAL
- `type` character; output binary type, one of: `Byte` for eight bit unsigned integer, `UInt16` for sixteen bit unsigned integer, `Int16` for sixteen bit signed integer, `UInt32` for thirty two bit unsigned integer, `Int32` for thirty two bit signed integer, `Float32` for thirty two bit floating point, `Float64` for sixty four bit floating point.
- `NA_value` non-NA value that should represent R’s NA value in the target raster file; if set to NA, it will be ignored.
- `update` logical; if TRUE, an existing file is being updated
- `normalize_path` logical; see `read_stars`
- `chunk_size` length two integer vector with the number of pixels (x, y) used in the read/write loop; see details.
- `progress` logical; if TRUE, a progress bar is shown
- `filename` character; used for guessing driver short name based on file extension; see examples
Details

write_stars first creates the target file, then updates it sequentially by writing blocks of chunk_size. in case obj is a multi-file stars_proxy object, all files are written as layers into the output file dsn

Examples

detect.driver("L7_ETMs.tif")

%in% stars-method

evaluate whether cube values are in a given set

Description

evaluate whether cube values are in a given set

Usage

## S4 method for signature 'stars'

x %in% table

Arguments

x data cube value
table values of the set
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